Kaasalainen, M.; Torppa, J.; Muinonen, K. (2001). "Optimization methods for asteroid lightcurve inversion. II. The complete inverse problem." *Icarus* **153**, 37-51.

Lagerkvist, C.I.; Magnusson, P. (1990). "Analysis of asteroid lightcurves. II. Phase curves in a generalized HG-system." *Astron. Astrophys. Suppl.* **86**, 119-165.

Slivan, S.M.; Binzel, R.P.; Boroumand, S.C.; Pan, M.W.; Simpson, C.M.; Tanabe, J.T.; Villastrigo, R.M.; Yen, L.L.; Ditteon, R.P.; Pray, D.P.; Stephens, R.D. (2008). "Rotation Rates in the Koronis Family, Complete to $H \approx 11.2$." *Icarus* **195**, 226-276.

Slivan, S.M. (2012). "Epoch Data in Sidereal Period Determination. I. Initial Constraint from Closest Epochs." *Minor Planet Bull.* **39**, 204-206.

Slivan, S.M. (2013). "Epoch data in sidereal period determination. II. Combining epochs from different apparitions." *Minor Planet Bull.* **40**, 45-48.

Slivan, S.M. (2014). "Sidereal Photometric Astrometry as Efficient Initial Search for Spin Vector." *Minor Planet Bull.* **41**, 282-284.

Slivan, S.M.; McLellan-Cassivi, C.; Shishido, R.; Wang, N. (2021). "Rotation Period of Koronis Family Member 1840 Hus." *Minor Planet Bull.* **48**, 112-113.

Tonry, J.L.; Denneau, L.; Heinze, A.N.; Stalder, B.; Smith, K.W.; Smartt, S.J.; Stubbs, C.W.; Weiland, H.J.; Rest, A. (2018). "ATLAS: A High-cadence All-sky Survey System." *PASP* **130**, 064505.

AN IMPROVED LIGHTCURVE AND ROTATION PERIOD OF 1178 IRMELA

Frederick Pilcher Organ Mesa Observatory (G50) 4438 Organ Mesa Loop Las Cruces, NM 88011 USA fpilcher35@gmail.com

Lorenzo Franco Balzaretto Observatory (A81), Rome, ITALY

Giulio Scarfi Iota Scorpii Observatory (K78), La Spezia, ITALY

Paolo Bacci, Martina Maestripieri GAMP - San Marcello Pistoiese (104), Pistoia, ITALY

Paolo Fini, Guido Betti Beato Ermanno Astronomical Observatory (L73) Impruneta, ITALY

Alessandro Marchini Astronomical Observatory, University of Siena (K54) Via Roma 56, 53100 - Siena, ITALY

(Received: 2023 Feb 26)

A collaboration of observers from North America and Europe has produced a lightcurve of 1178 Irmela with synodic period 11.992 ± 0.001 hours, amplitude 0.29 ± 0.02 magnitudes, and full phase coverage.

Previously published period determinations for 1178 Irmela are reported by Binzel (1987), 19.17 hours; Stephens (2012), 11.989 hours; and Polakis (2019), 11.985 hours. The lightcurves published by Stephens (2012) and by Polakis (2019) were made from a single observatory and therefore show only about 2/3 phase coverage for an Earth-commensurate period. Stephens (2012) showed that the 19.17-hour period published by Binzel (1987), based upon sparse lightcurves widely spaced in time, is an 8/5 alias of a period near 11.99 hours.

The authors of this paper, widely spaced in longitude in North America and Europe, respectively, agreed to collaborate to obtain full phase coverage. An equipment list for all observers is provided in table II. Ten sessions 2023 Jan. 26 - Feb. 17 provide a very good fit to bimodal lightcurve with synodic period 11.992 ± 0.001 hours and amplitude 0.29 ± 0.02 magnitudes with full phase coverage. This result is very close to the periods published by Stephens (2012) and by Polakis (2019), and to 5/8 of the period by Binzel (1987).

On 2023 Jan. 31, P. Bacci and M. Maestripieri obtained alternating data points in the V and R filters that show V-R = 0.40. This value is within the usual range 0.38 ± 0.05 for asteroids with C-type taxonomic classifications (Shevchenko and Lupishko, 1998).

Number Name	yyyy/mm/dd	Phase	LPAB	Врав	Period(h)	P.E	Amp	A.E.	
1178 Irmela	2023/01/26-02/17	* 7.4, 5.8	138	-5	11.992	0.001	0.29	0.02	

Table I. Observing circumstances and results. The phase angle is given for the first and last date, unless a minimum (second value) was reached. LPAB and BPAB are the approximate phase angle bisector longitude and latitude at mid-date range (see Harris et al., 1984).

Observatory (MPC code)	Telescope	ССД	Filter
Organ Mesa Observatory (G50)	0.35-m SCT f/10	SBIG STL-1001E	С
Iota Scorpii(K78)	0.40-m RCT f/8	SBIG STXL-6303e (bin 2×2)	Rc
GAMP (104)	0.60-m NRT f/4	Apogee Alta	V,Rc
Beato Ermanno Astronomical Observatory (L73)	0.30-m SCT f/6	QHY174M CMOS (bin 2×2)	Rc
Astronomical Observatory, University of Siena (K54)	0.30-m MCT f/5.6	SBIG STL-6303e (bin 2×2)	С

Table II. Observing Instrumentations. MCT: Maksutov-Cassegrain, NRT: Newtonian Reflector, RCT: Ritchey-Chretien, SCT: Schmidt-Cassegrain.





References

Binzel, R.P. (1987). "A Photoelectric Survey of 130 Asteroids." *Icarus* 72, 135-208.

Harris, A.W.; Young, J.W.; Scaltriti, F.; Zappala, V. (1984). "Lightcurves and phase relations of the asteroids 82 Alkmene and 444 Gyptis." *Icarus* **57**, 251-258.

Polakis, T. (2019). "Photometric observations of seventeen minor planets." *Minor Planet Bull.* **46**, 400-406.

Shevchenko, V.G.; Lupishko, D.F. (1998). "Optical properties of Asteroids from Photometric Data." *Solar System Research* **32**, 220-232.

Stephens, R. (2012). "Asteroids observed from GMARS and Santana Observatories: 2011 October - December." *Minor Planet Bull.* **39**, 80-82.